



# BIST Founding Conference

Growing together, advancing science



Barcelona Institute of  
Science and Technology

## Elementary Particle Detectors and Data Handling Systems: A historical perspective

Prof. Manuel Delfino, UAB and IFAE, Director of PIC

Dialogues on the boundaries of knowledge: Big Data

MARCH 31, 2017

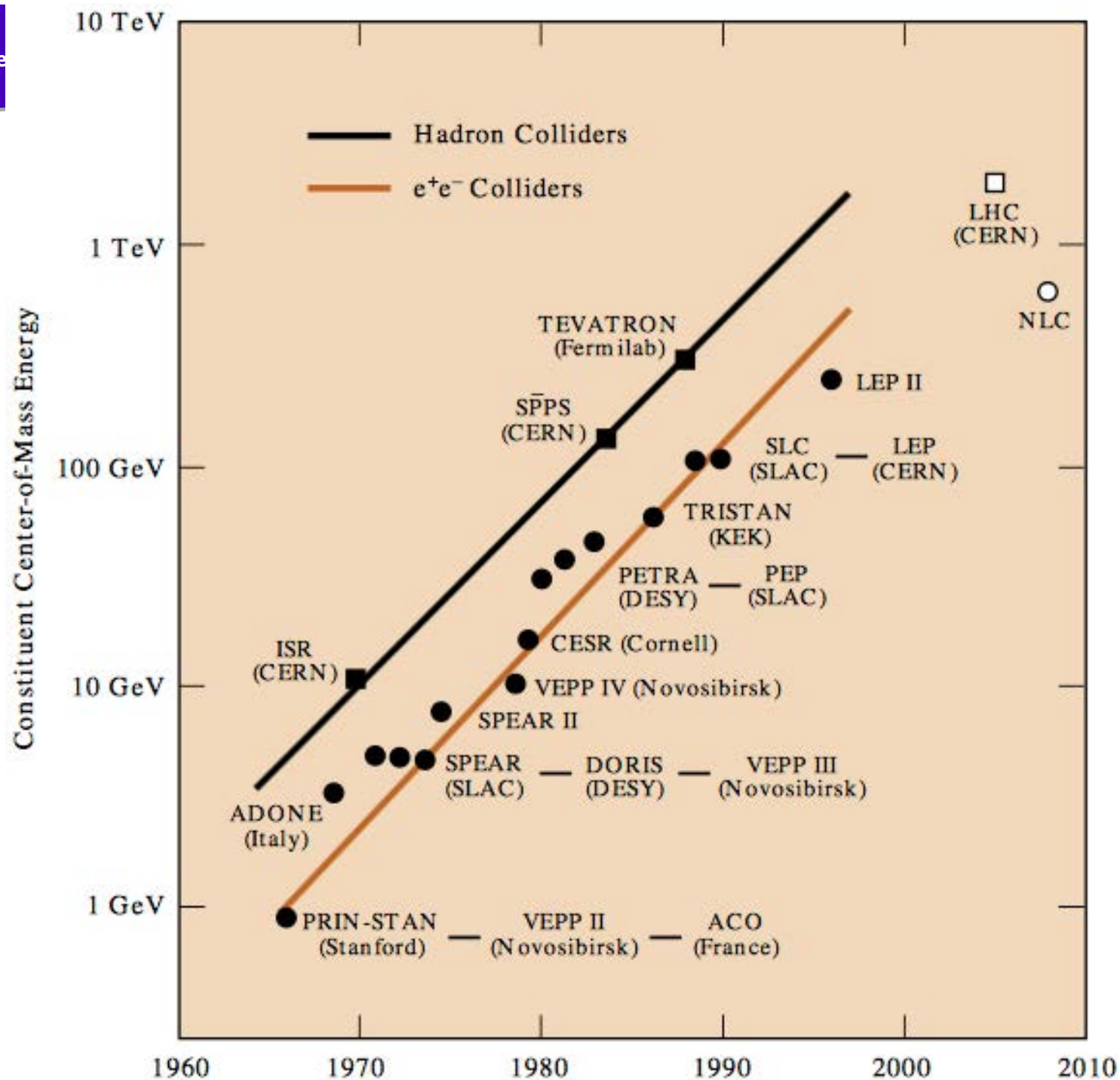
CCCB | Barcelona

## Why Experimental Particle Physics = High Energy ?

- Particle accelerator beam  $\rightarrow$  Particle mass + Kinetic Energy
- Collision: Beam Energy  $\rightarrow$  (many, different) particles + Kinetic Energy
  - Many possible outcomes  $\rightarrow$  final state topology to differentiate
  - Energy-momentum conservation  $\rightarrow$  Initial Kinetic Energy limits the maximum mass that can be produced in a collision

**If an accelerator doesn't find what you are looking for, the only way out is to move on to a higher energy one**

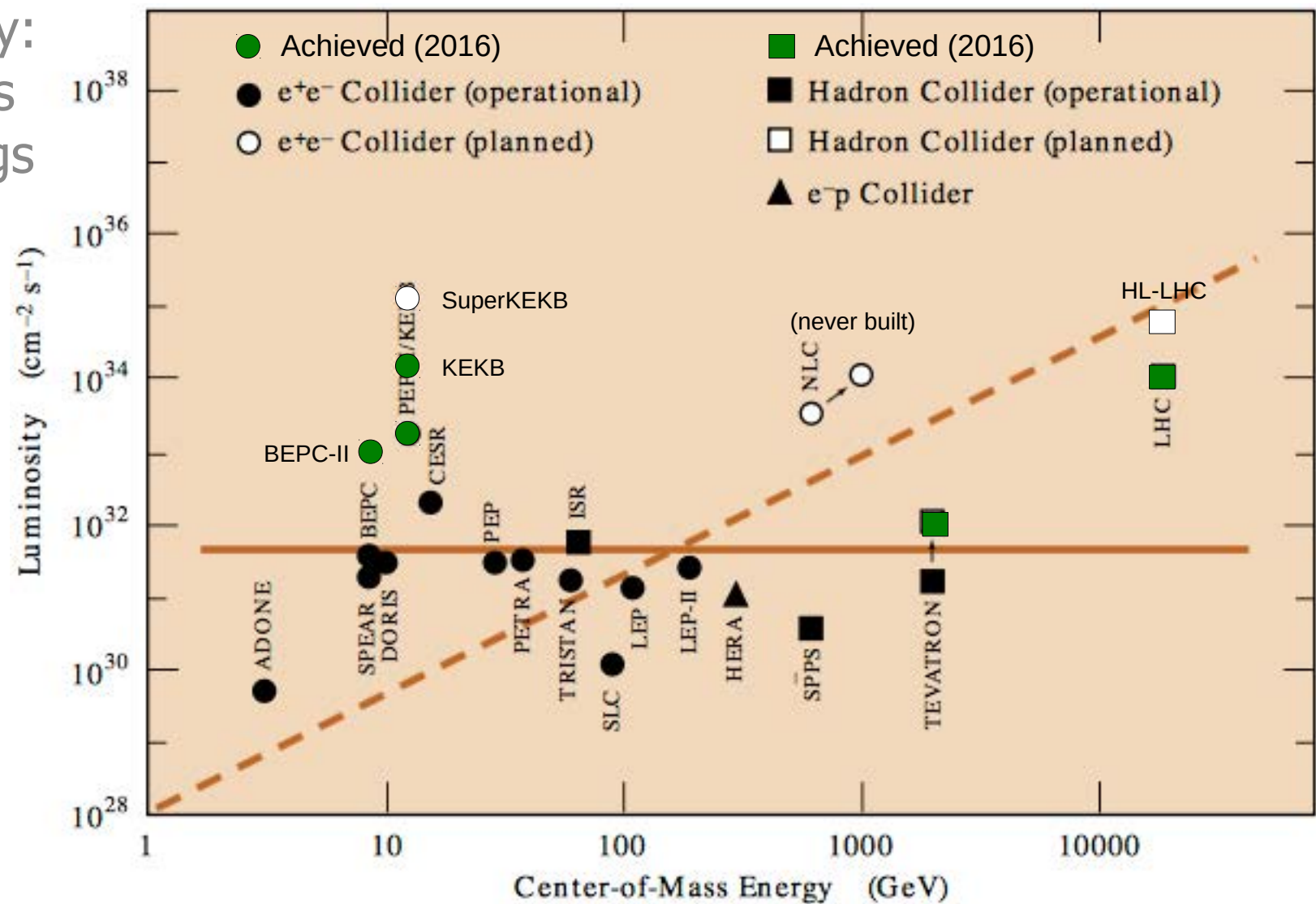
**...except: If the probability of producing it is too small, strive to make your accelerator more "luminous"**



Panofsky,  
SLAC Beam Line,  
Spring 1997,  
Vol. 27, No. 1

## More luminosity:

- More particles
- More crossings
- Tighter focus



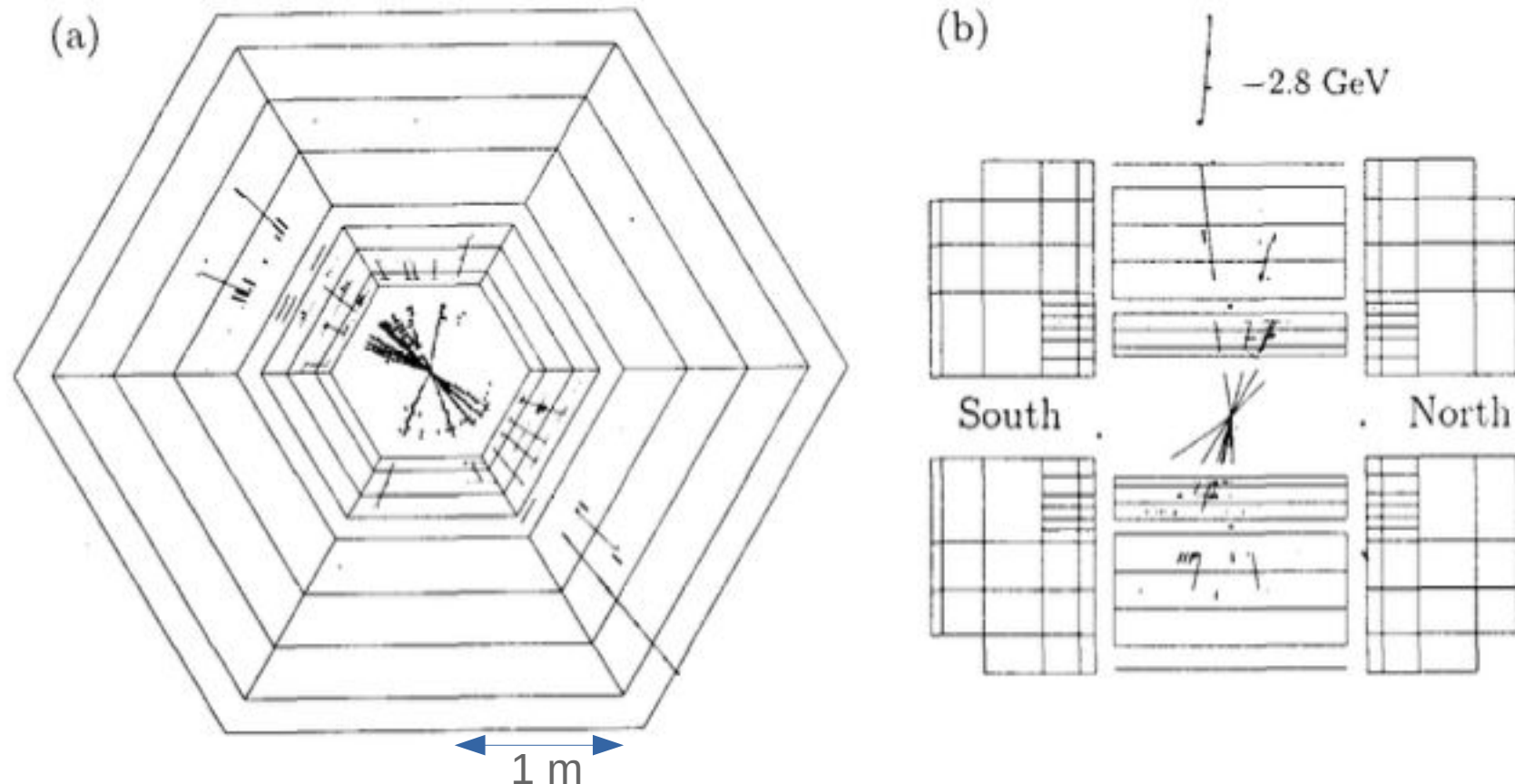
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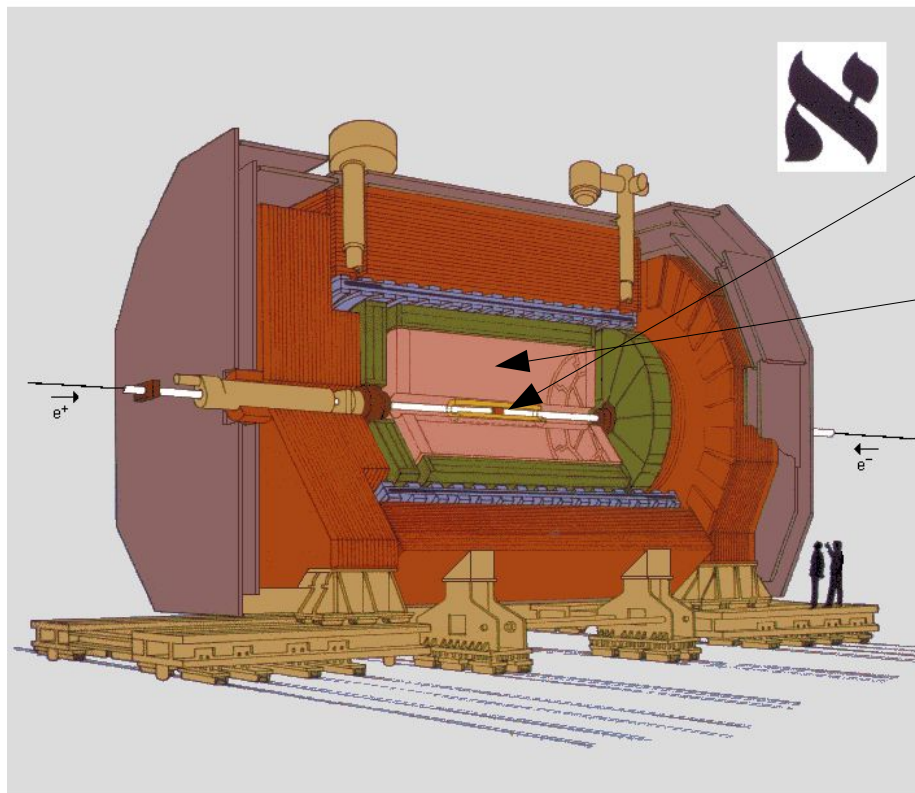
Note: Neutrino factories not shown









## Detectors: Larger, more granular, faster, smarter

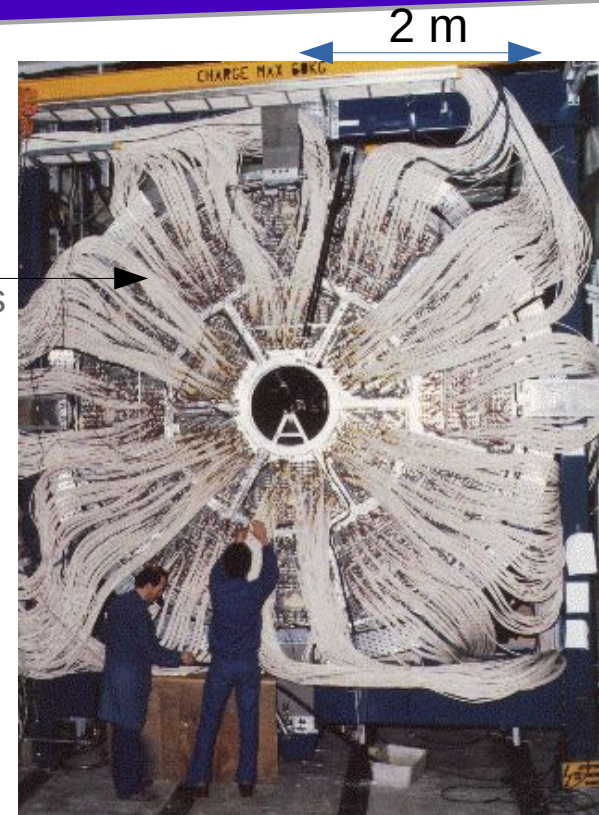
- Use bending in magnet to measure speed of produced particles
  - Magnetic force = Centripetal force  $q v B = \frac{m v^2}{R}$   $v = \frac{q}{m} B R$
  - B limited → Need increasing precision in trajectories
- Number of particles produced in a collision increases with energy
  - Particles emerge in “jets”, very crowded together
  - High-mass particles are unstable and decay shortly after produced
  - Need to identify the types of particles
- More luminosity:
  - 400 kHz in 1980's → 40 MHz in LHC today
  - 1 collision in 10 crossings 1970's → 20 collisions in 1 LHC crossing
- More online filtering of more complex topologies.  
Yesterday's interesting physics becomes today's noise.



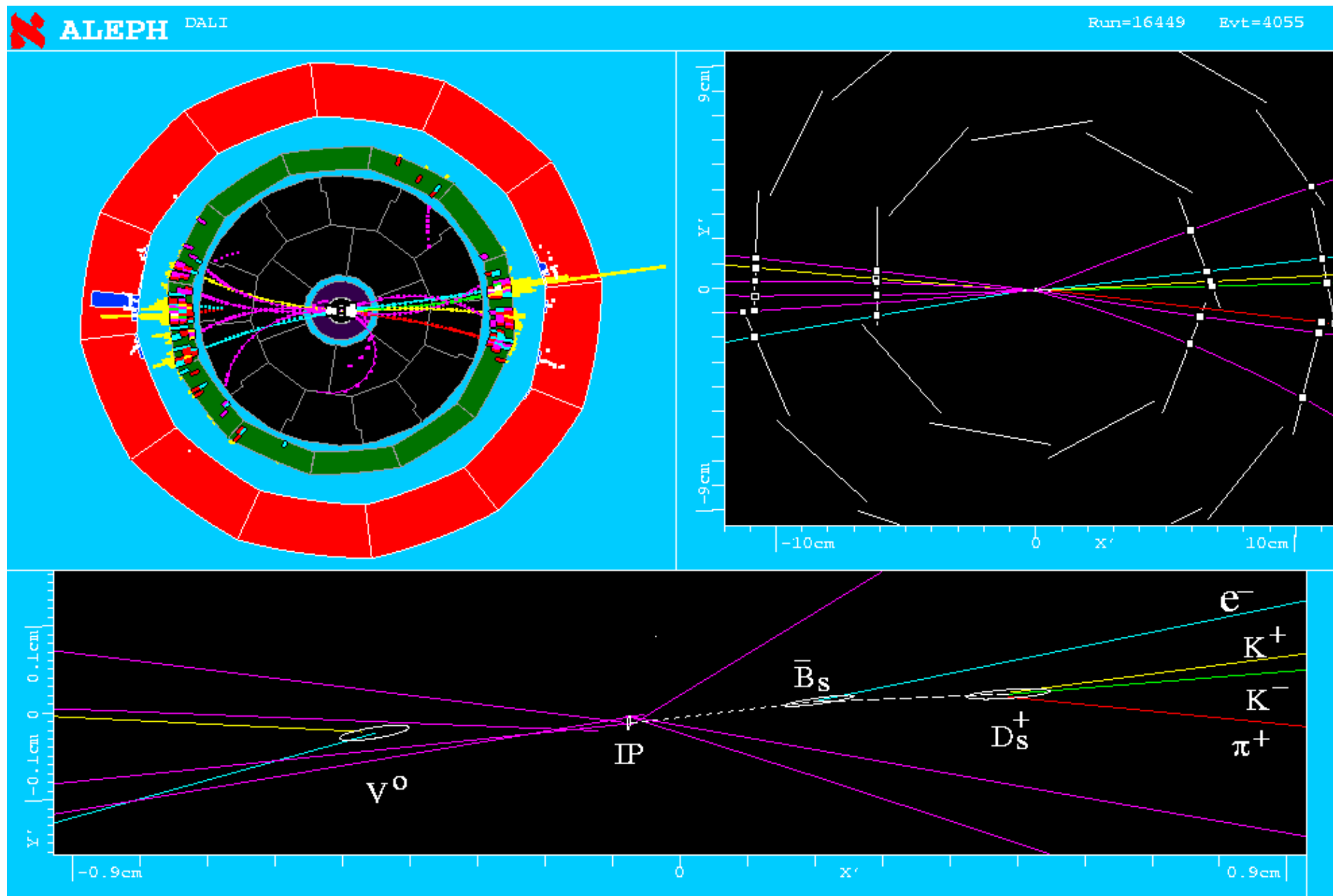
- Digitization electronics managed by Motorola 68k microprocessors
- Data acquisition + event filter: VAX minicomputer (4 MB RAM, 1 GB disk, 170 MB/tape)
- 0.5 Mbps ~mile link to IBM mainframe → remote data recording, automated workflows
- Analysis supported by a *single* IBM mainframe for the *whole* laboratory – Virtual Machines
- No data access from participating institutions

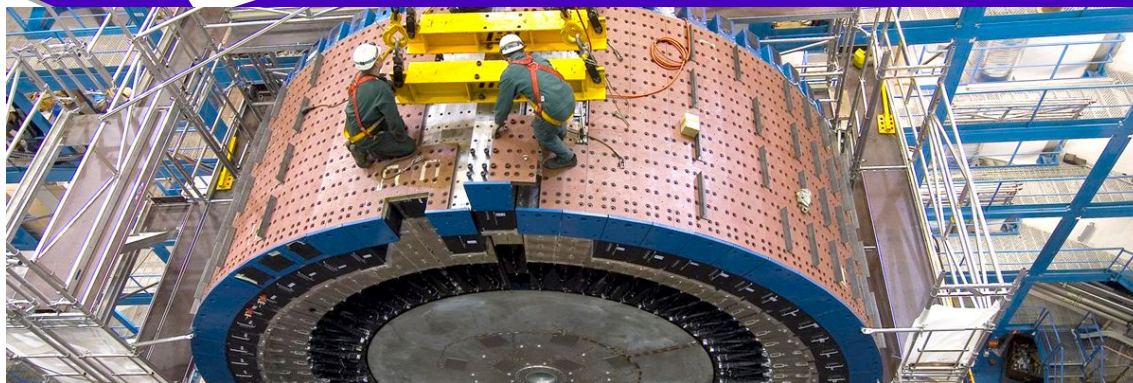


-  Vertex Detector 100k channels
-  Inner Tracking Chamber
-  Time Projection Chamber 50k channels
-  Electromagnetic Calorimeter
-  Superconducting Magnet Coil
-  Hadron Calorimeter
-  Muon Chambers
-  Luminosity Monitors



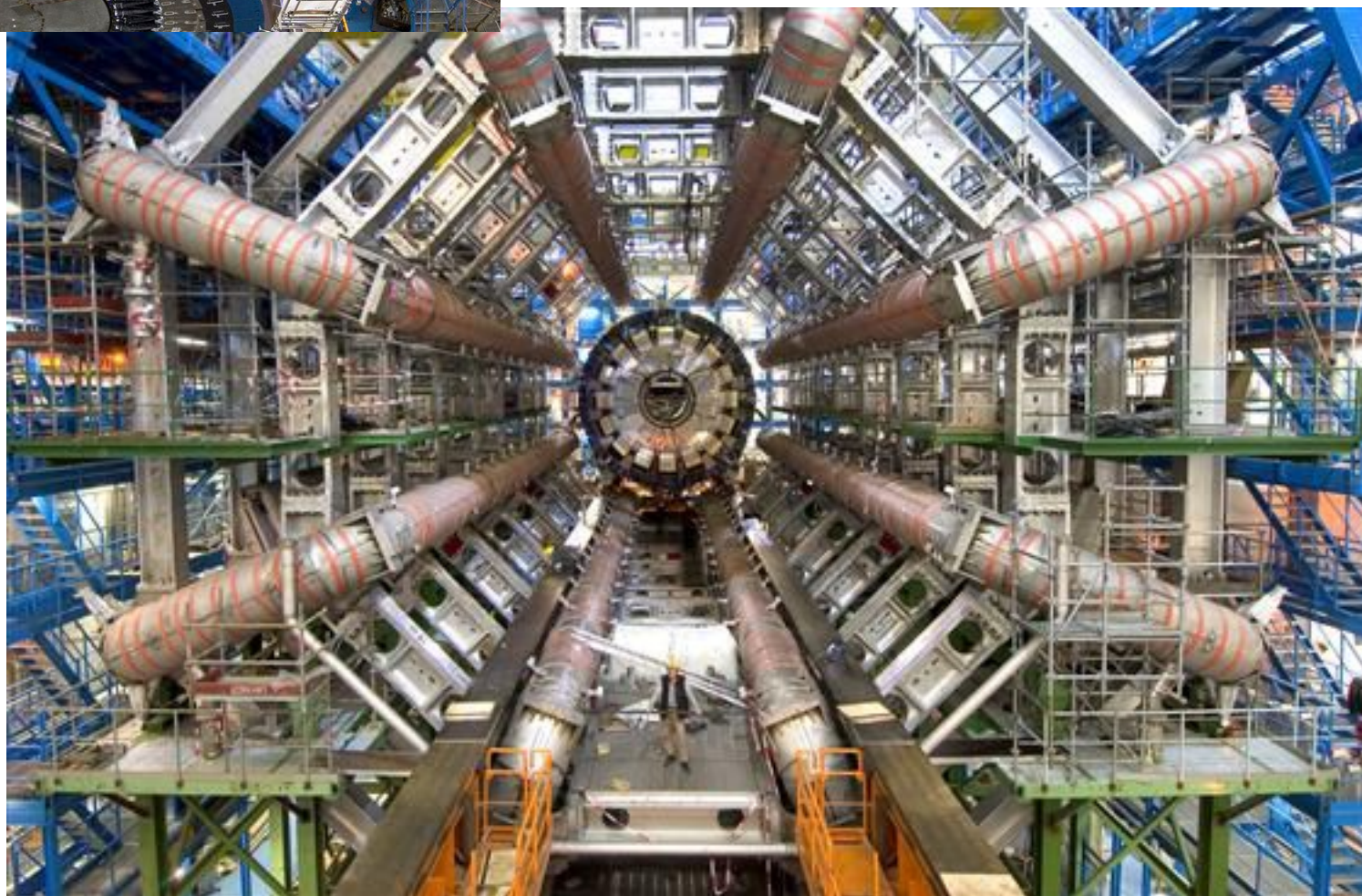
- Precision timing over 1000 m<sup>3</sup> volume. Flash ADC's, microprocessors, software trigger
- Dedicated detector-side data processing farm (originally 12 VAX workstations on Ethernet)
- CERN fiber optic network: remote data recording to CERN data center (>5 km away)
- Data analysis:
  - Shared IBM mainframe complemented by private VAX workstation cluster
  - Duplicate data tapes shipped to some laboratories (mainly France and UK)
  - Later overtaken by SHIFT workstation and PC farms developed and operated by CERN IT

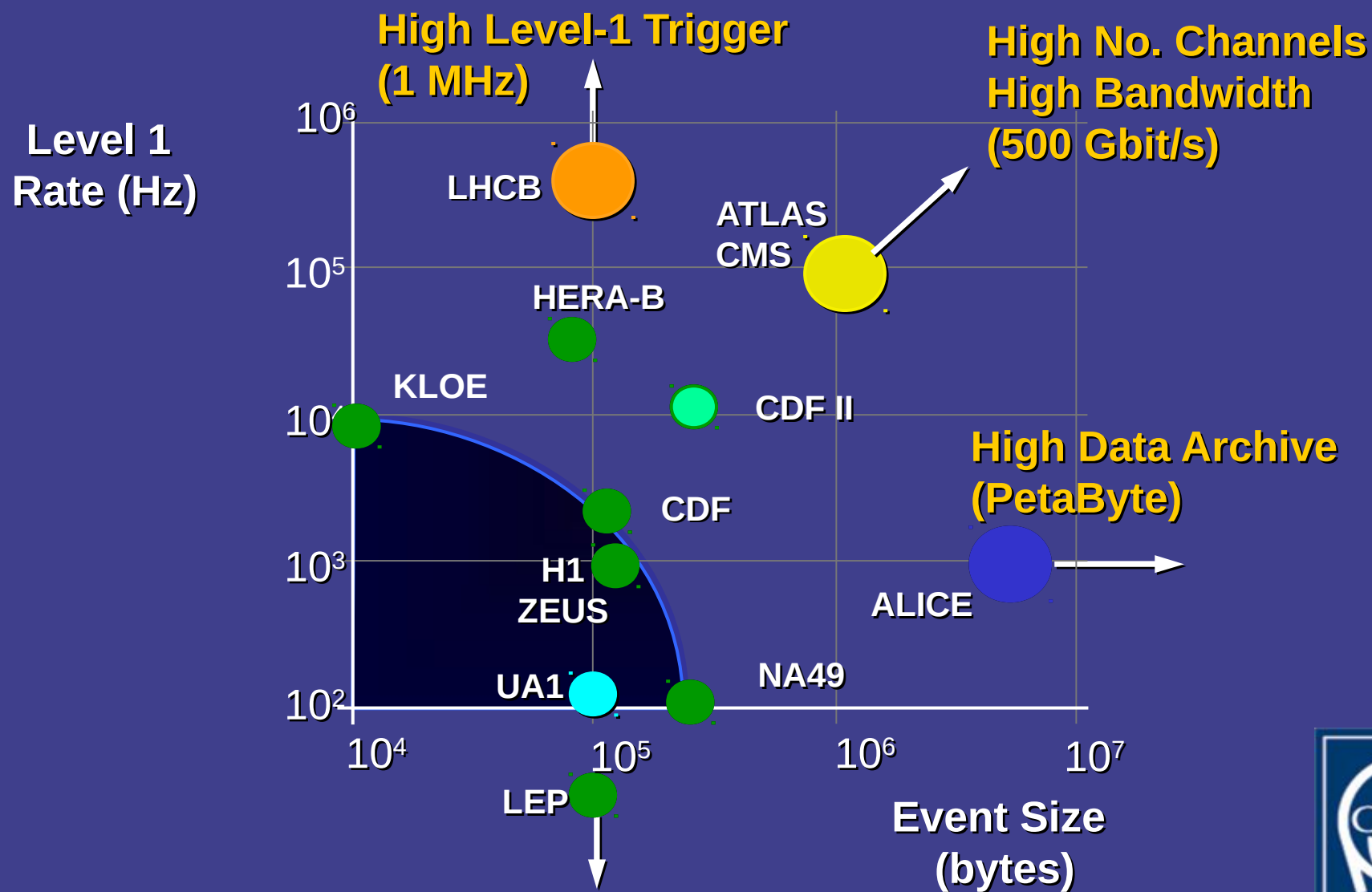


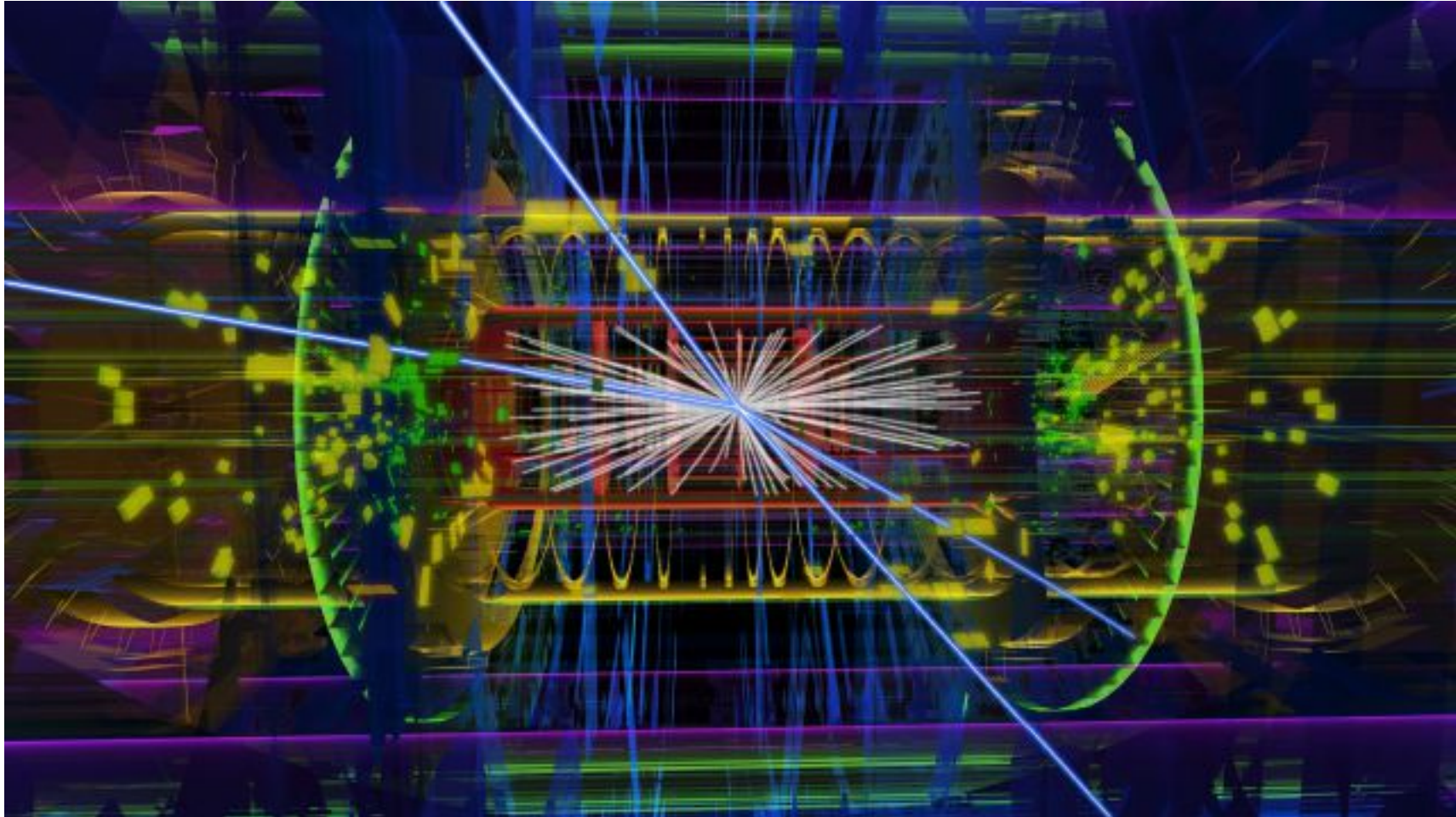


- 46 m long  
25 m diameter  
7000 Tons
- Inner track detector:  
80M pixels  
6M silicon strips  
350k channel TRT
- Data acquisition system  
handles 60 Petabytes  
per second
- Trigger+Event Filter  
select 100 Hz out of 1  
GHz ( $1 \text{ in } 10^7$ )
- More than 20 TB per  
day of recorded data

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## **Astrophysics and Experimental Cosmology**

- Big Bang → Early Universe dominated by elementary particle physics
- Standard Model of Particles: Quantum Gravity is a big missing piece
- Expansion of the Universe is accelerating, don't understand why

## **New high-tech telescope-like instruments have become an important complement to accelerators**

- Air Cherenkov gamma-ray “telescopes”
- Instrumented water or ice neutrino “telescopes”
- Laser interferometer gravitational wave detectors
- Sky surveys: ground-based or satellite-born automated telescopes
- Synthetic aperture radio telescopes

## **Significant computing and storage needs with complex management, access and preservation issues**



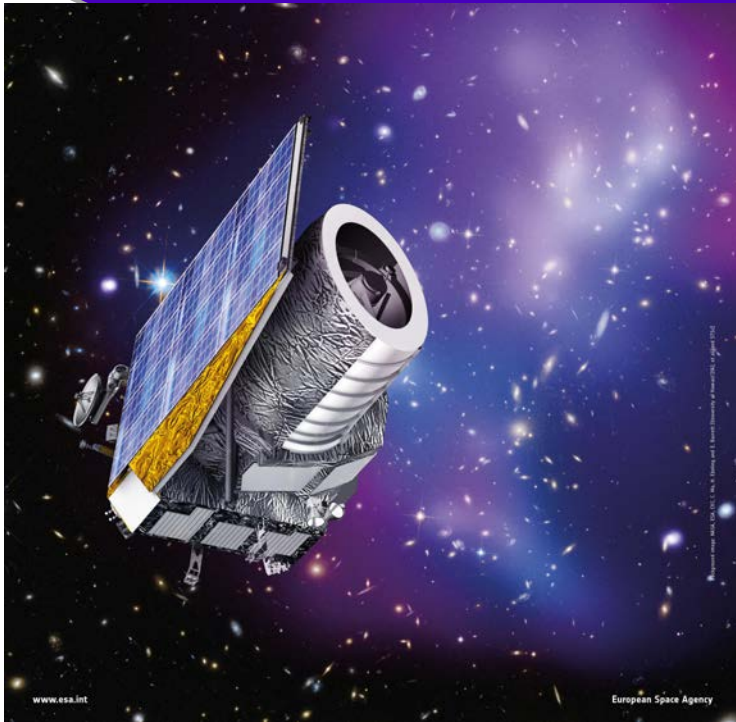
Artist rendering of Northern Array at ORM, La Palma, Canary Islands



1,350 members from 210 institutes in 32 countries



Large Telescope concrete pad at ORM



Artist rendering of EUCLID at the L2 Lagrangian point

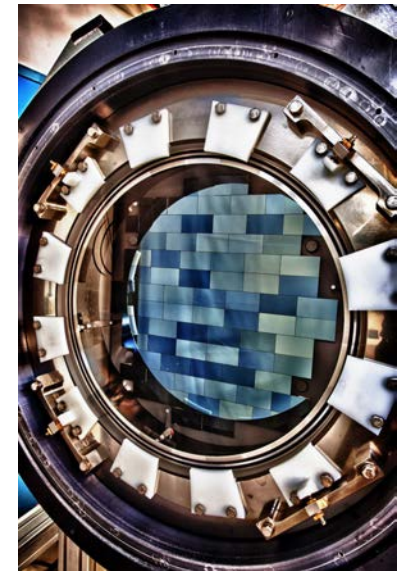
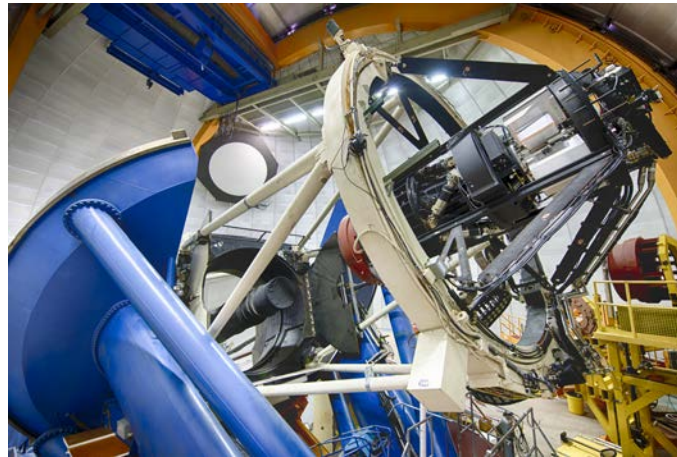
- 1400 members from 180 institutes in 16 countries
- 10 billion sources
- Data in the tens of Petabytes, possible by improved down-link speed from space to Earth
- Science Data Centers (similar to LHC Tier-1s)



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- 400 scientists from 25 institutions in 7 countries
- Measuring 300 million galaxies
- Data management at U.S. National Center for Supercomputing Applications (NCSA) in Illinois
- PIC “Big Data” platform to speed-up analysis

Dark Energy Camera (DECam) on Blanco Telescope in Chile (Reidar Hahn, Fermilab)



DECam Focal Plane Detector (Reidar Hahn, Fermilab)



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**Over the last 50 years, scientists and engineers have worked hand in hand to build bigger and better instruments for particle physics, astrophysics and experimental cosmology.**

**These instruments are generating data of increasing complexity in increasing volumes, which must be shared at a planetary scale.**

**Innovation in computing and data processing domain is a key activity ... and Frédéric Hemmer from CERN will tell you more about it !**



# PIC

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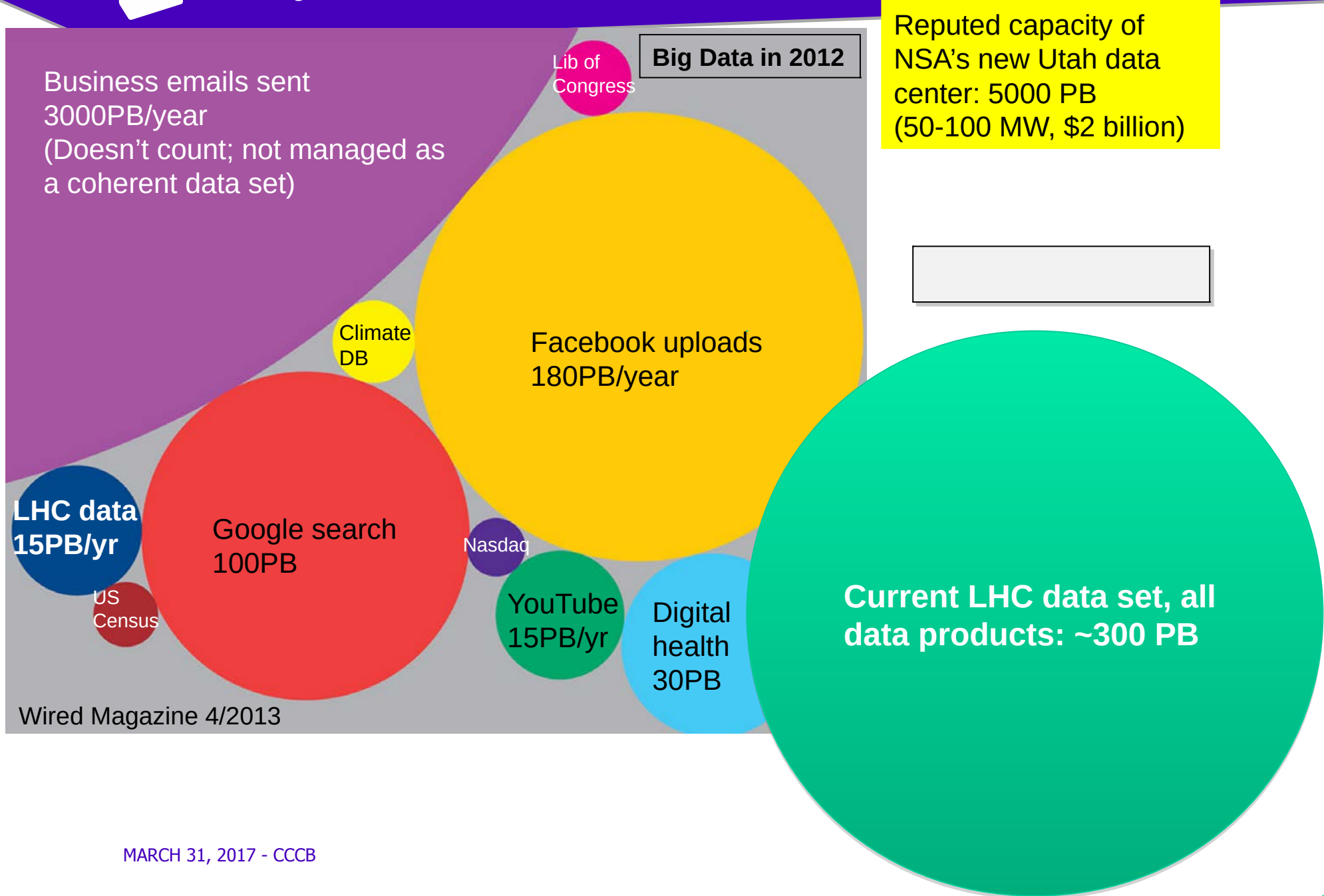


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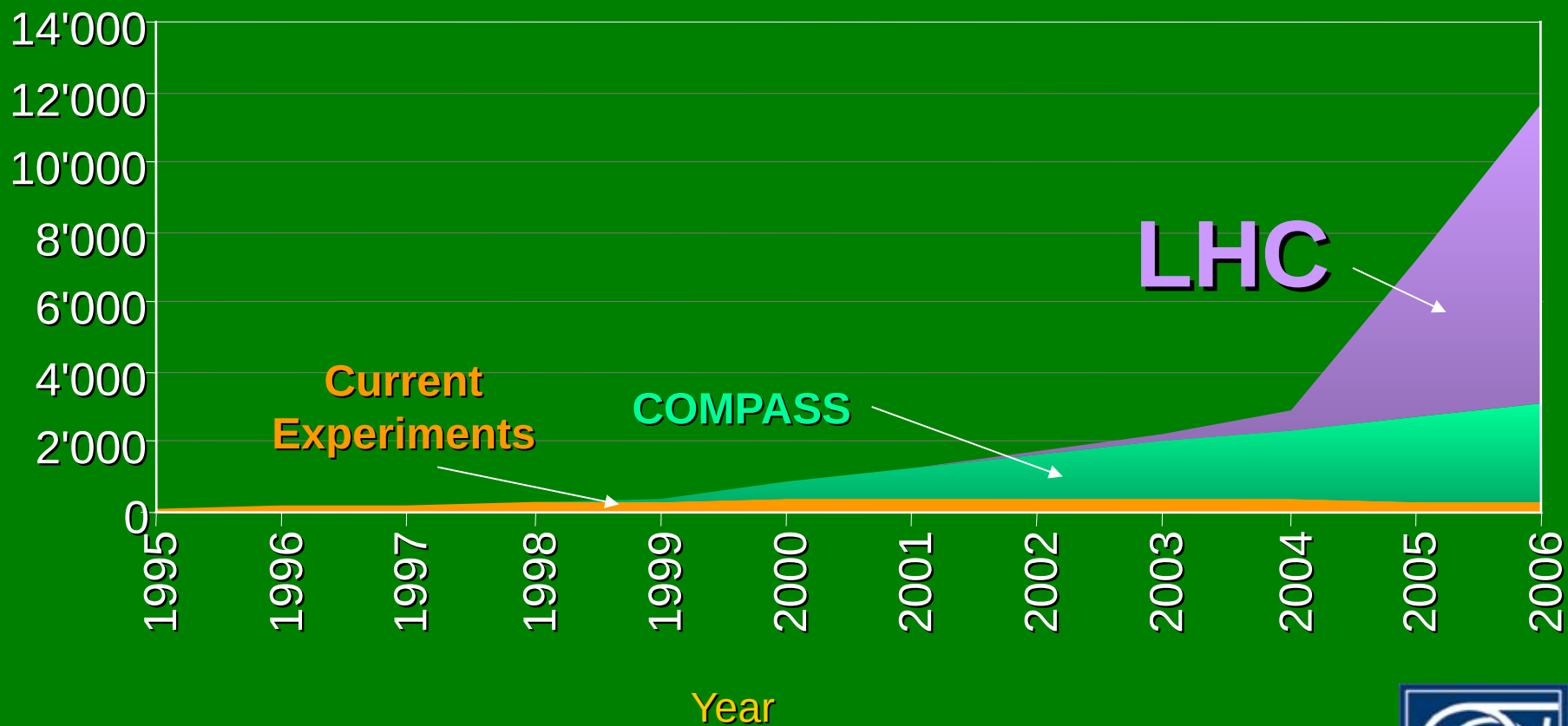
# Additional slides



LHC Will Accumulate Multi-Petabytes

TeraBytes

Long Term Tape Storage Estimates



Complex Data = More CPU Per Byte

## Evolution of Computing Capacity

